Lattice-Boltzmann Models of Ion Thrusters

Dr. Jacques C. Richard

richard@aero.tamu.edu

And

Prerit Shah

Prerit_aggie@neo.tamu.edu

Texas A & M University

College Station, TX 77845, USA



Lattice-Boltzmann Method (LBM) & Ion Thrusters

- Complement Discrete Simulation Monte-Carlo (DSMC) models for faster computation of critical ion thruster parameters
- Use LBM to model plasma flow in thruster
- Compare results with experimental data and DSMC predictions
- Identify plasma flow characteristics that lead to thruster component erosion; *e.g.*, grids



Outline

- Ion thrusters (see Gallimore, 2004)
 - Basic physics of operation
 - Issues of interest: lifetime/erosion
- Why try LBM?
- LBM & Ion thrusters
- Some results
- Summary, conclusion & future work



Ion Thrusters



Ion thrusters are the most efficient EP devices at converting input power to thrust and are used both as primary propulsion and for station-keeping on commercial and scientific spacecraft.

Key issues include grid erosion and thrust density limitations from space-charge effects.

Ion Thrusters Basics



Ion thruster concept (Gallimore, 2004)

- Electrons are emitted from discharge cathode assembly (DCA)
- DCA electrons (*Primary*) are accelerated by local sheath to high voltage (>15 eV)
- *Primary* electrons create ions via impact ionization with neutrals
- Ionization process starts with one *Primary* and one neutral results in 2 *Maxwellian* electrons and one ion
- Ions are attracted to ion optics (Screen grid) via electric field
 - Ions are accelerated through optics (Screen & Accel grids) - ion beam neutralized by neutralizer cathode
 - Accel grid negative to prevent electron backstreaming
 - Note: While *Maxwellian* electrons outnumber *Primaries* 10:1, the latter account for most of the ionization in the discharge chamber.



Modern Ion Thrusters

Solar Electric Propulsion — NASA's Evolutionary Xenon Thruster (NEXT) [5-10 yr. deployment time] -NEXT is the follow-on to NSTAR used on DS1 and slated for DAWN (2006 launch)

-NEXT represents a 4x improvement in thrust and power and a 25% increase in Isp (from 3280 to 4100 s) over NSTAR at half the specific mass (from 2.6 to Nuclear Electric Propulsion — NASA's Nuclear Space Initiative [10-15 yr. deployment time] Electric Propulsion Proposals in NASA's 2002 "In-Space Propulsion Technologies" NASA Research Announcement (NRA) for ultra-high-performance engines (Isp > 6,000 s)









NASA





Ion Thruster Basics



Typical Ion Engine Parameters

 Within a few cm of grid, typical ion thruster & plasma parameters are:

 $n_{Xe+} \sim 10^{12} \text{--} 10^{10} \text{ cm}^{-3}$, $n_e \sim 10^{12} \text{--} 10^{10} \text{ cm}^{-3} >> n_{Xe} >> n_{Xe++}$...

- V₊ ~ 1075 V at screen grid
- V₋ ~ -150 V at accelerator grid
- Grid separation ~ 1 mm
- Screen grid opening diameter ~ 2 mm
- Accelerator grid opening diameter ~ 1 mm



Assumptions of Applicability of LB

Note local *Kn*: $Kn = \frac{\lambda}{L} = \frac{RT}{\sqrt{2}\pi d^2 N_A pL}$ Crawford (2002) $L = \frac{\rho}{d\rho / dx}$

- $Kn \sim O(0.1)$ around optics
- Ion veloc. distrib. Laserinduced Fluorescence
 Velocimetry of Xe II in the 30-cm NSTAR-type Ion
 Engine Plume, Smith and Gallimore (AIAA-2004-3963)
- Maxwellian radial f(v)



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LBM EP Model

The model assumes coupling of the velocity distribution function w/the electrostatics

$$\frac{\partial f}{\partial t} + \mathbf{c} \bullet \nabla_{\mathbf{r}} f - \frac{q}{m} \nabla_{\mathbf{r}} \phi \bullet \nabla_{\mathbf{c}} f = Q(f, f)$$
$$\varepsilon_0 \nabla_{\mathbf{r}}^2 \phi = e \int f d^3 \mathbf{c}$$

- Assume linear collision if close to the continuum limit so that $Q = -v_r(f f^{(eq)})$
- Adequate for near equilibrium plasma, simple charge exchange (CEX) collisions or even assume "Q=0" for collisionless, electricallydriven plasma



Computational Procedure





Axi-symmetric Cylindrical Coordinates

- In accordance w/thruster geometry
- Use work of Yu, Girimaji & Yu (2004) where cyl. coord. effects are incorporated via source terms in LBE to satisfy macrolevel cyl. coord. eqs. (NS)

$$g_{\alpha} = w_{\alpha}s + \frac{3}{c^2}w_{\alpha}\mathbf{e}_{\alpha}\cdot\mathbf{a}$$

where,

$$s = -\frac{u_r}{r}, a_z = \frac{v}{r}\frac{\partial u_z}{\partial r} + \frac{q}{m}\frac{\partial \phi}{\partial z}\frac{\partial f}{\partial v_z}, a_r = \frac{v}{r}\left(\frac{\partial u_r}{\partial r} - \frac{u_r}{r}\right) + \frac{q}{m}\frac{\partial \phi}{\partial r}\frac{\partial f}{\partial v_r}$$



Results

- Compare general trends: nondimensional
- Compare specific cases



LBM Ion Thruster Exhaust Stream

Unitless ion #density contours; matches Crawford (2001)





LBM Ion Thruster Exhaust Stream

Ion velocity field





LBM Ion Thruster Exhaust Stream

Electrostatic potential





 To look at grid erosion, we want to zoom in on a grid segment with 2D/axi-symmetric models as below





Electrostatic potential contours from modeling a slit btwn grids





LBM Ion Thruster Optics

Electrostatic potential





 Zoom in on a optics segment with 2D/axisymmetric models as below





Electrostatic potential btwn grids





Electrostatic potential contours from Gallimore (2004)





LBM Ion Thruster Optics





Ion # density; screen grid 10V, accelerator grid -10





Ion # density; screen grid 1075V, accelerator grid -180





Ion velocity field; screen grid 1075V, accelerator grid -180





Conclusions & Future Work

- LBM does well w/modeling EP
- Next is to try
 - other species
 - variations in collision operator, e.g.,pseudo-random collision frequency as used in DSMC
 - Other variations of BE form



Computational Domain



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Extrapolation Boundary

- Extrapolation boundary condition has been applied at S1 and S4 in computational domain.
- ≻ At S1,
 - f(1,j,9) = f(2,j,9)
 - f(1,j,2) = f(2,j,2)
 - f(1,j,6) = f(2,j,6)
- ➤ At S4,
 - f(NX,j,8) = f(NX-1,j,8)
 - f(NX,j,4) = f(NX-1,j,4)
 - f(NX,j,7) = f(NX-1,j,7)





Free stream boundary

- Free stream boundary condition has been applied at S2 and S3 in the computational domain.
- ➢ At S2,
 - f(1,j,9) = 0.0
 - f(1,j,2) = 0.0
 - f(1,j,6) = 0.0
- ➤ At S3,
 - f(NX,j,8) = 0.0
 - f(NX,j,4) = 0.0
 - f(NX, j, 7) = 0.0





Symmetric Boundary

Symmetric boundary condition has been applied at S5 on the computational domain.

≻ At S5,

- f(i,1,7) = f(i,2,8)
- f(i,1,3) = f(i,2,5)
- f(i,1,6) = f(i,2,9)



